

Implications of  $m_H \sim 125 \text{ GeV}$

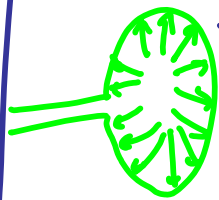
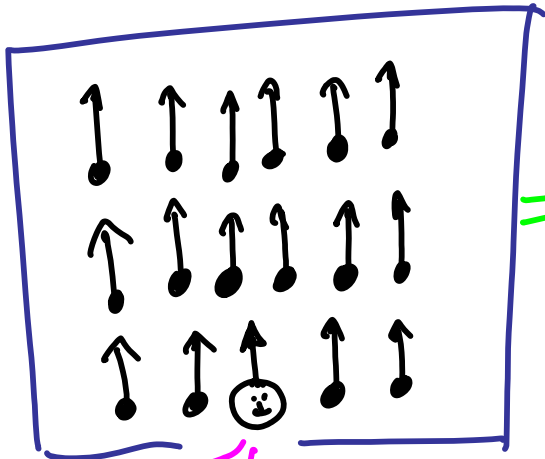
# TRIUMPH OF WEAK COUPLING

TECHNICOLOR

1978 - 2011

R.I.P.

Never seen before in "state of nature"

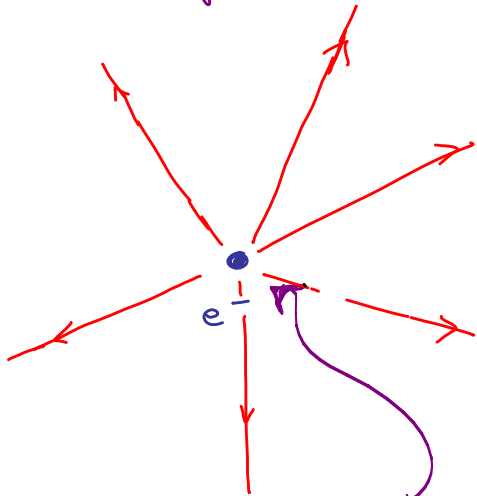


"fine-tuning"

Why are we all pointed in same direction?

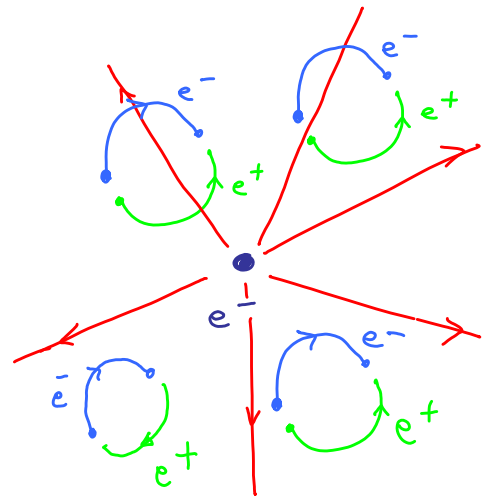


# Infinite Energy In Electric Field



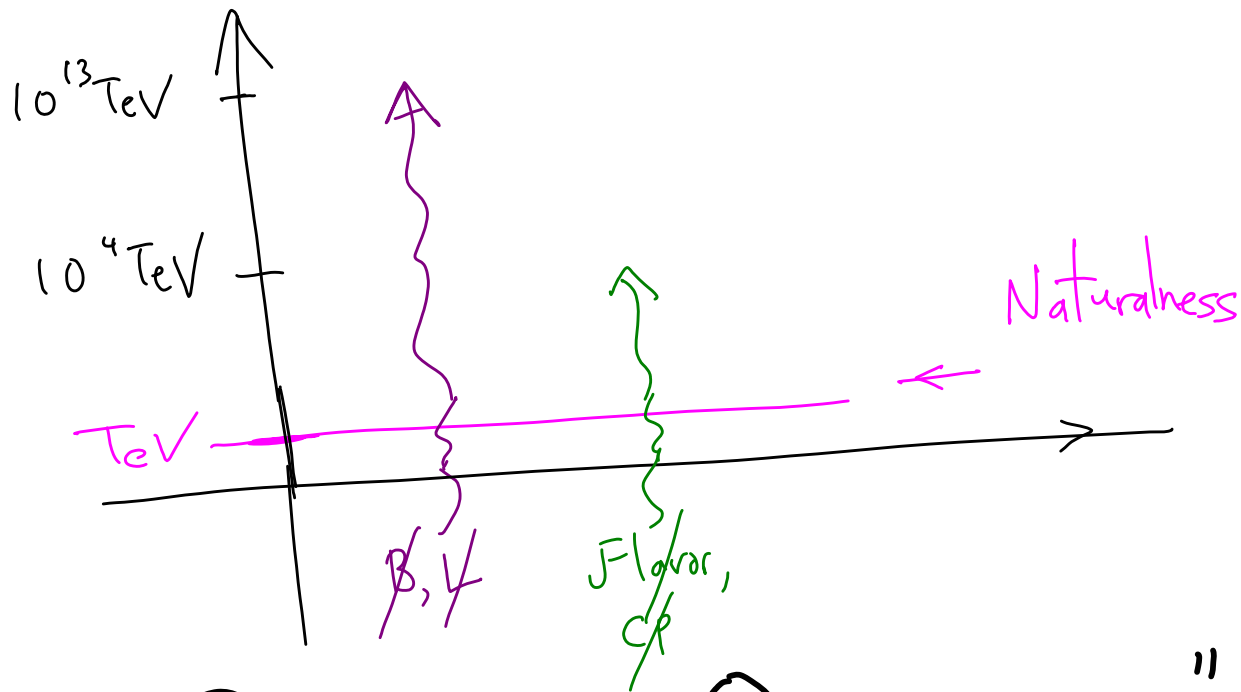
$$\frac{e^2}{4\pi a_{cl}} \sim m_e c^2 \Rightarrow a_{cl} \sim \frac{e^2}{4\pi m_e c^2}$$

Q.M.  
Rel.



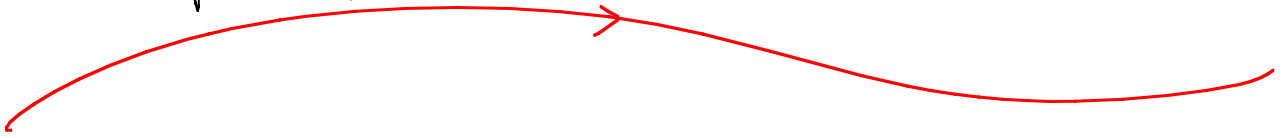
$$a_{\text{Compton}} \sim \frac{\hbar}{m_e c}$$
$$\sim \left[ \frac{e^2}{4\pi \hbar c} \right]^{-1} \times a_{cl} !$$

# Tension Driving BSM Physics For 30 yrs

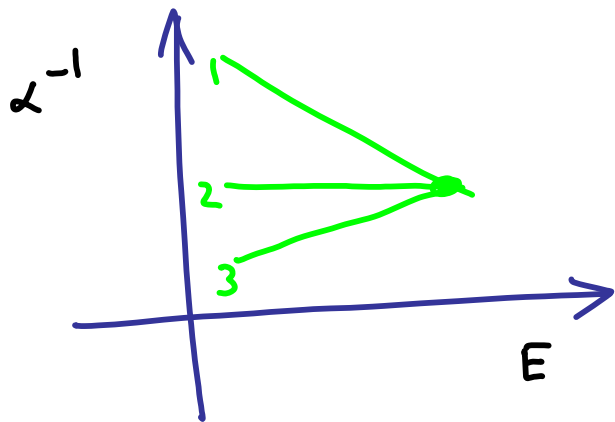


" NOT PROBLEMS - OPPORTUNITIES "

Natural SUSY



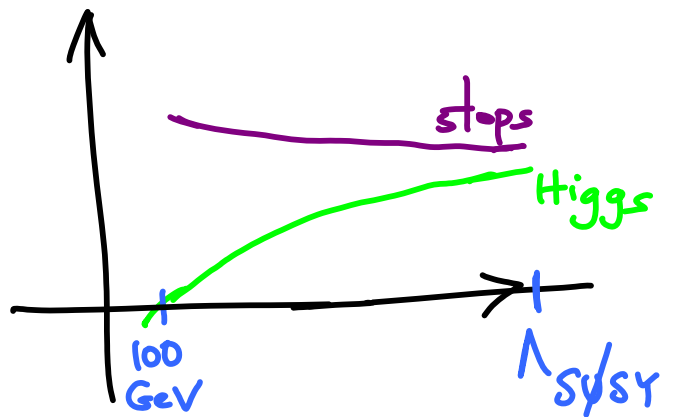
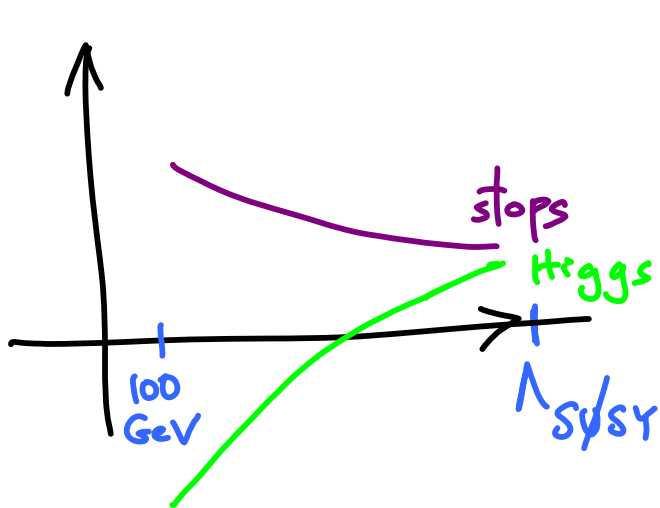
SUSY circa 1990



+ D.M.

SPECTACULAR

# Why No SUSY @ LEP?



—  $Z, h, \tilde{t}, \tilde{g} \dots$   
 —  $\tilde{l}, EW \text{ Kinus}$   
 "Natural" Spectrum

?  
 ↑  
 —  $Z, h$   
 "Nature-al" spectrum

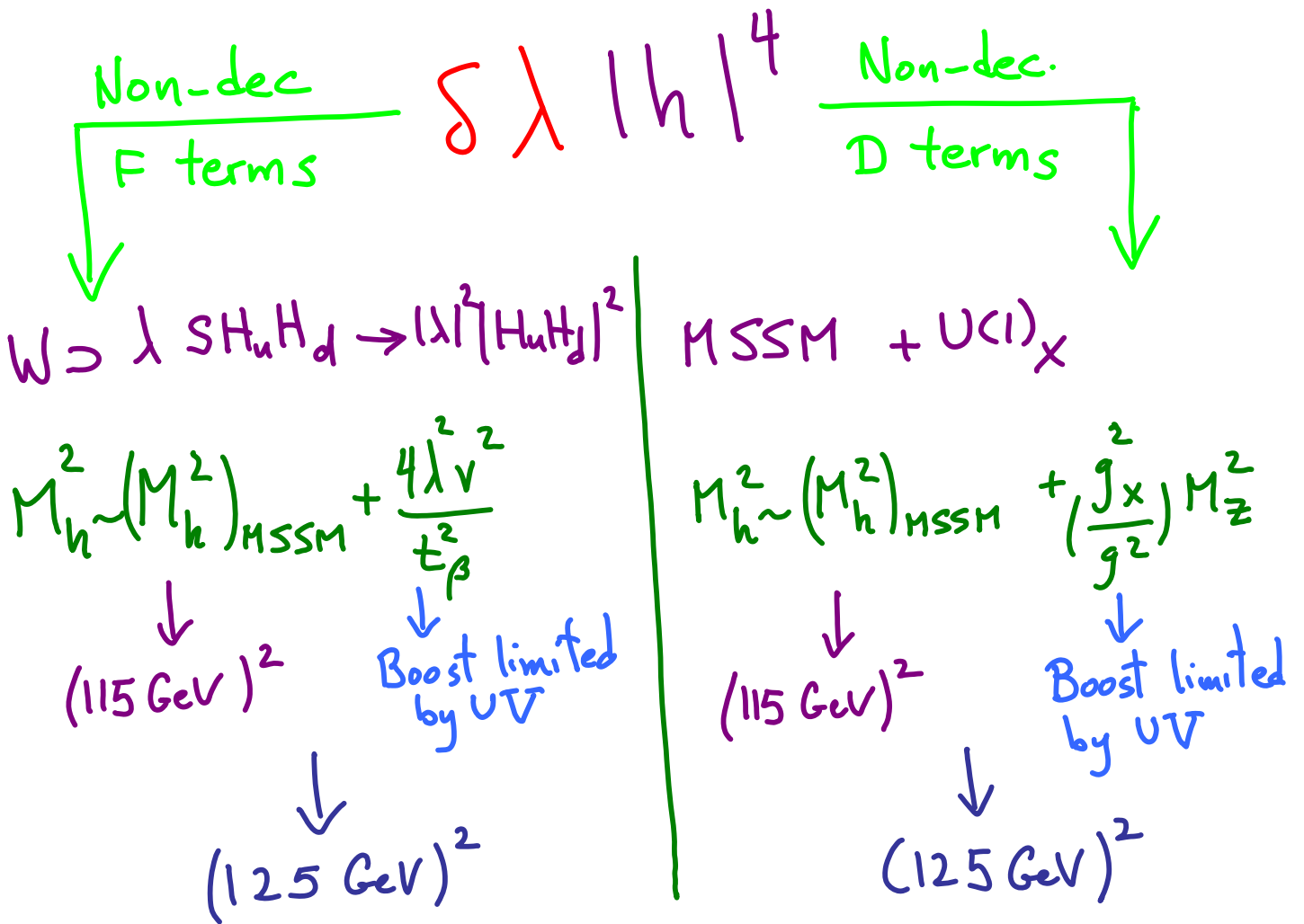


• In the MSSM, a milder cousin of this problem is

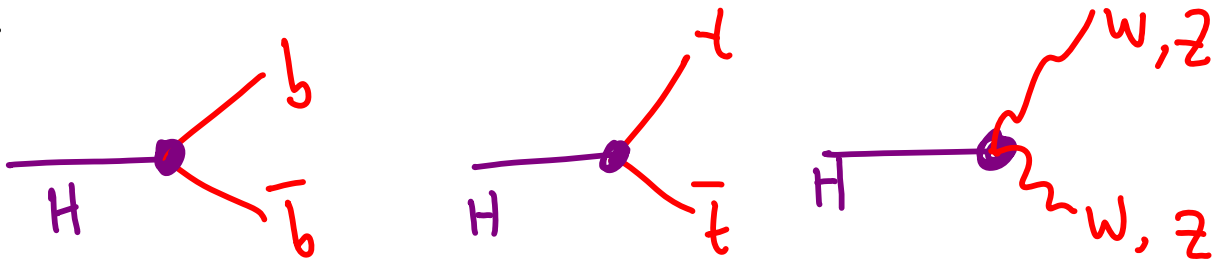
$$M_h \lesssim M_Z + M_h^{\text{sys, stops}}$$

↑  
threshold +  
log-enhanced

With  $m_h \sim 125$ , this alone is a  $\sim 1\%$  tuning, @ best.



{ A ~ 20% deviation in



→ must have new, extra light states  
important for EWSB, aside from Higgs  
→ HUGE support for Natural Theories!

Note:  $m_H \sim 125 \text{ GeV}$  is pretty

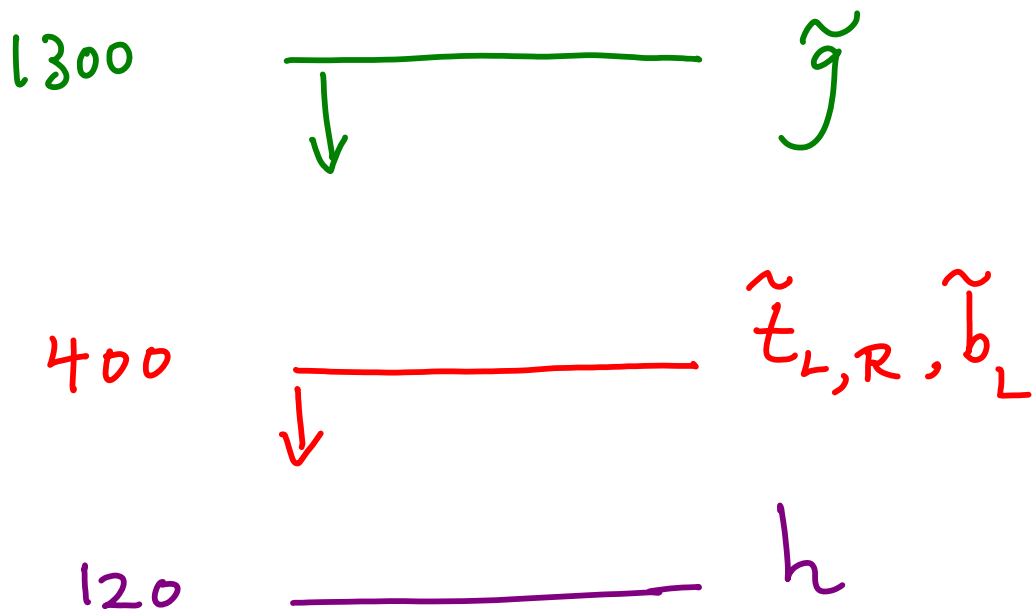
close to  $m_Z$ ! SUSY is only

theory where this proximity is

reasonable, since  $\lambda = g^2$ . { For Higgs

as PGB,  $\lambda_{\text{top}}$  sets potential,  $m_H \sim 250 \text{ GeV}$   
most reasonable }

# Compulsory Natural SUSY



Unavoidable tunings:  $\left(\frac{400}{m_{\tilde{t}}}\right)^2$ ,  $\left(\frac{4m_{\tilde{t}}}{M_{\tilde{g}}}\right)^2$

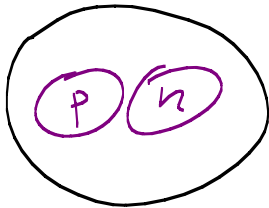
Minimally Split SUSY

or

"Don't Be Too Clever!"

# Could the weak scale be tuned?

Next EFT down:



De binding energy  
 $\sim 2 \text{ MeV}$

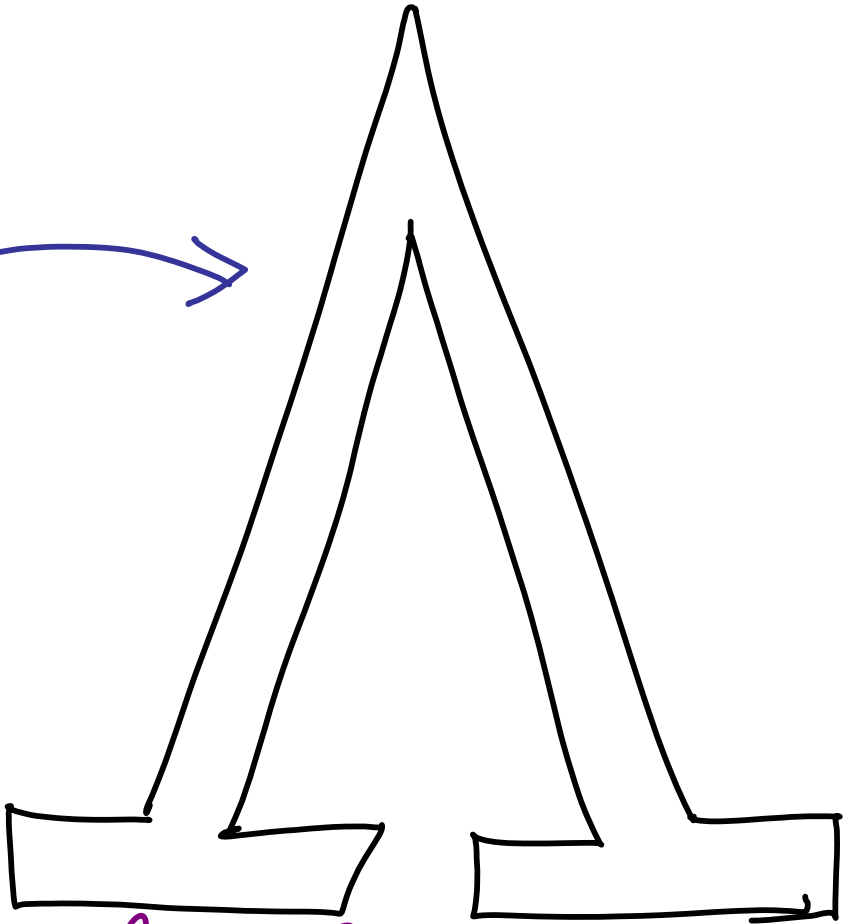
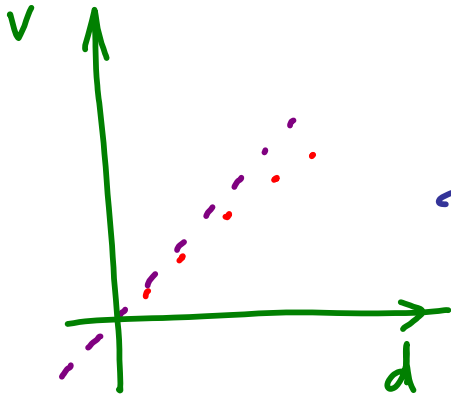
1 in 20 tuning

$\sim 10^{-4}$  tuning



Not bound by  
 $\sim 60 \text{ keV} (!!)$

few in 1000 tuning



NATURALNESS



Plausible anthropic tuning for weak  
scale: existence of atoms other than  
hydrogen



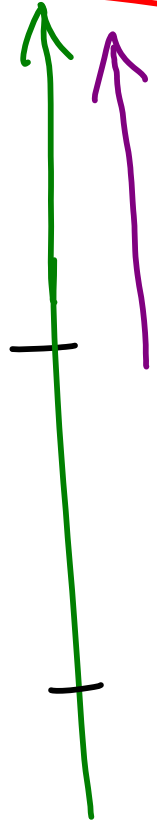
$(m_n - m_p) \gtrsim$  nuclear binding energies

# SPLIT SUSY

Reason for splitting:  
fermions carry R-symmetry,  
scalars don't

100's  
TeV

TeV



Scalars

Fermions

Unification ✓

Dark Matter ✓

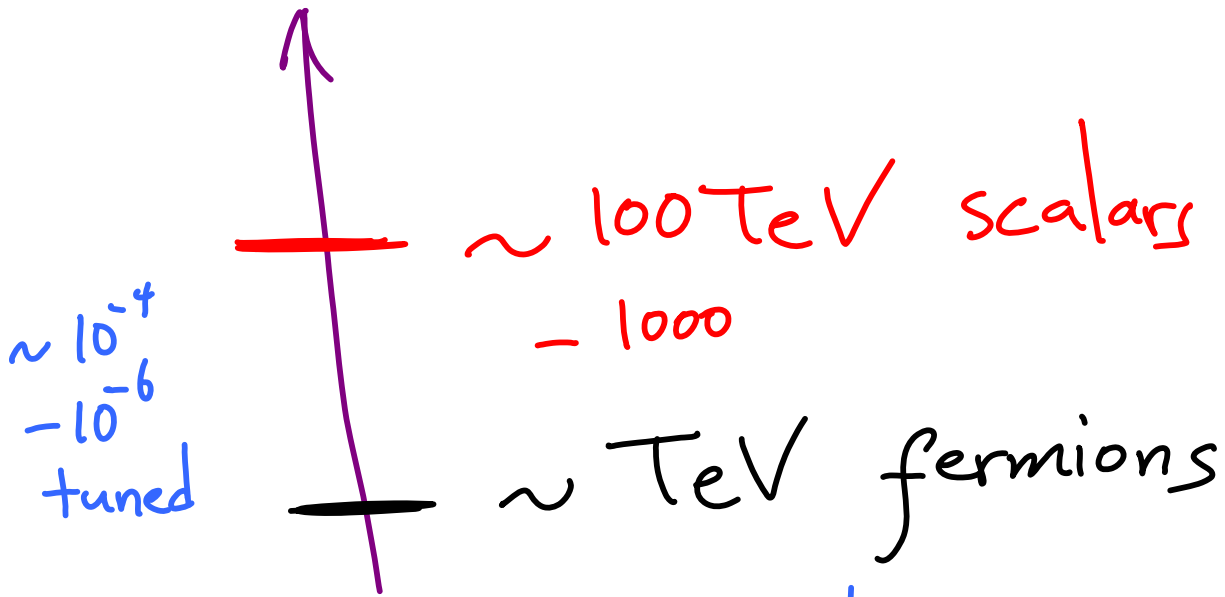
NO Flavor,  
CP, moduli, ...  
problems

Indeed, a 1-loop splitting between  
gauginos + scalars was ubiquitous  
in simplest models of ~~SUSY~~.

Modern guise: H.S. ~~SUSY~~: Anomaly med.  
Unless you work!  
["Not problem -  
opportunity"]

$$m_\lambda \sim \frac{\alpha}{4\pi} m_{3/2}$$
$$m_S \sim m_{3/2}$$

# THIS SPECTRUM



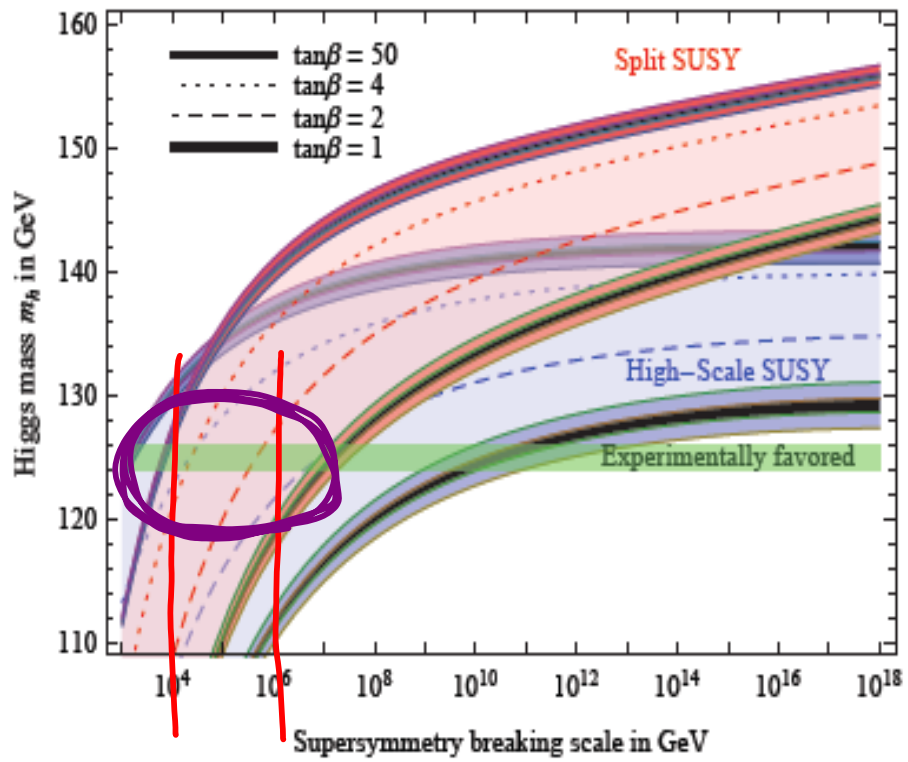
IS WHAT SUSY MODELS  
"WANT TO DO". LET THEM!

# Simplest Split SUSY

Also  $\mu \sim m_S \sim m_{3/2}$  [Giudice-Masiero]

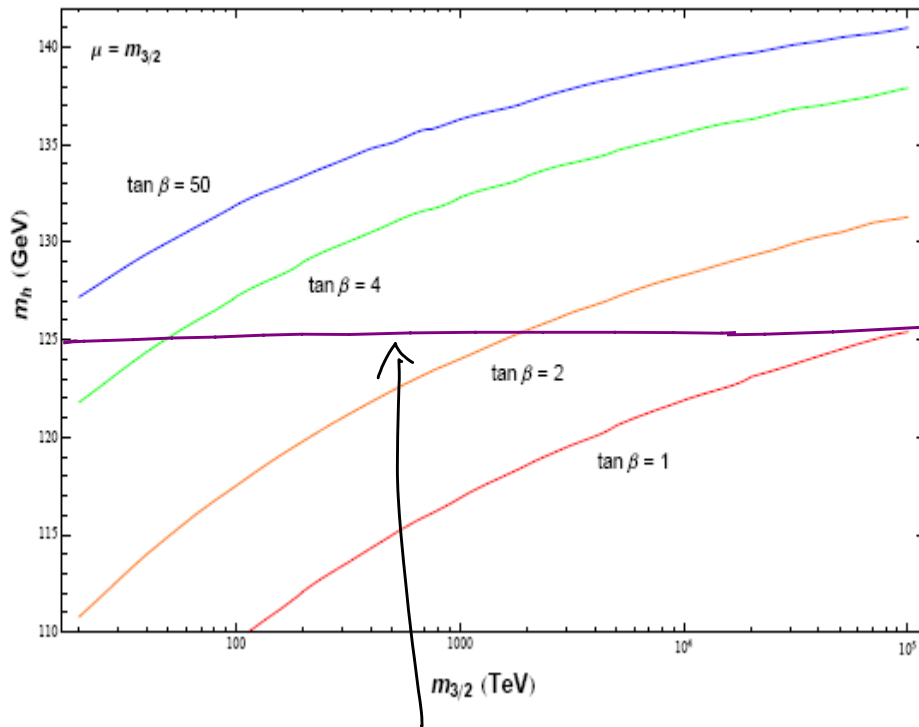
So only gluino, wino, bino light

### Predicted range for the Higgs mass



[ Giudice, Strumia ]

# Simplest Split SUSY



$\sim 1000$  TeV stops,  $\tan \beta \sim 3$  perfectly ok.  
[Mindless Solution of Flavor Problem]

$$M_h \sim 125 \text{ GeV}$$

Points to the most moderate "split" spectrum - with 10 - 1000 TeV scalars - which is also what SUSY "want to do".  
[NOT most radical story with  $m_s \gg 10^3 \text{ TeV}$ ]



# An Idea With a Long History

Early 80's : 1-loop split spectrum common in SUSY models

1999 : Giudice-Luty-Murayama-Rattazzi anom-med.  
(not Randall-Sundrum version, no sequestering!)

Split Spectrum thought of as  
embarrassment

2003 : J. Wells "PeV-scale SUSY"

2004 : NAH + Savas, Giudice + Romanino "Split SUSY"

2006 : NAH, Delgado, Giudice "Simplest Split SUSY"

⋮ Many others

# The Obvious Worry

"WIMP miracle" relic could mean

$m_{\text{Higgsino}} \sim 1 \text{ TeV}$  or  $m_{\text{Wino}} \sim 3 \text{ TeV}$ .

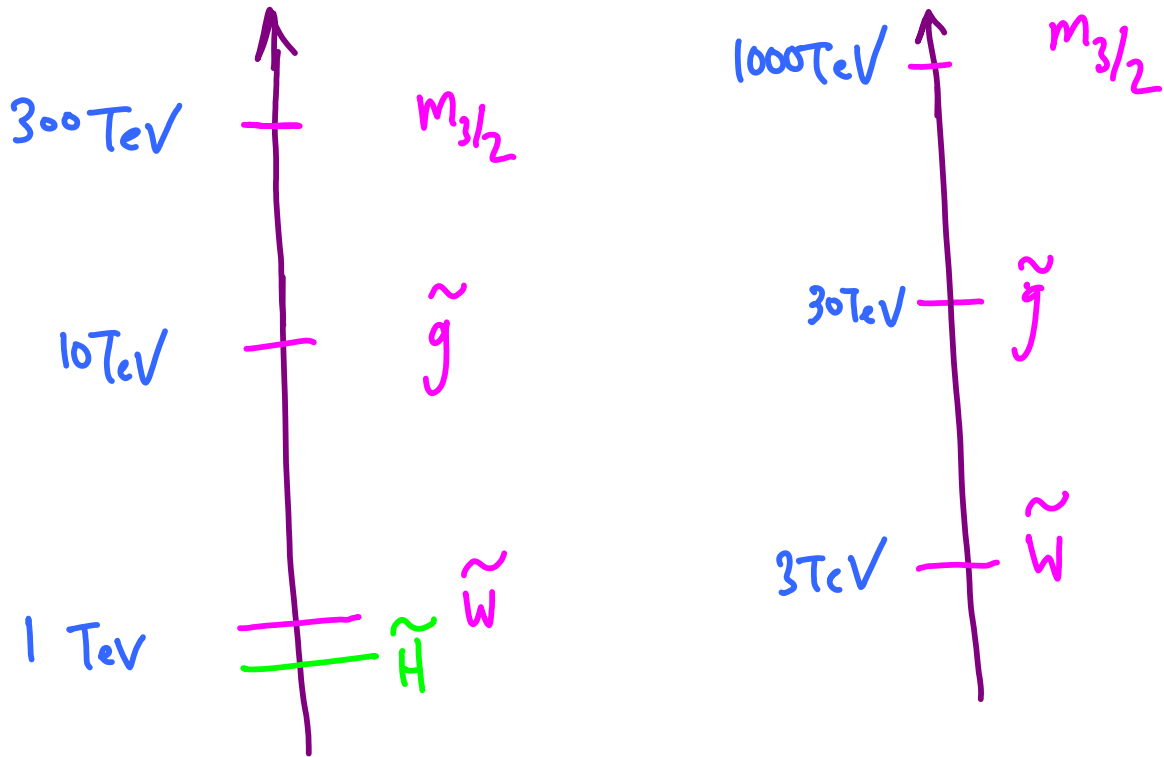
Won't ever see even fermionic  
superpartners.

Ⓐ Could still be "well-tempered" relic,  
lighter OK.

Ⓑ Relic could constitute part of DM [rest,  
say, axions]  $\rightarrow$  lower bottom of spectrum.

Ⓒ Moduli near  $\sim 100$  TeV decay late, dilute  
initial abundance [+ largely solve axion overclosure problem!],  
repopulate DM still favors light  $\sim$  TeV LSP.

# "Standard" Anomaly Mediation



## SUSY Vector-Like Masses

$$\int d^4\theta \phi^\dagger \phi (F^\dagger F + F^{\dagger c} F^c)$$

$$+ \int d^2\theta \phi^3 M F F^c$$

$\Downarrow$

$$\int d^4\theta F_{can}^\dagger F_{can} + \int d^2\theta (\phi M) F_{can}^c F_{can}$$

$\left(\frac{B\mu}{M}\right) = +m_{3/2}$ , keep on anom-med trajectory

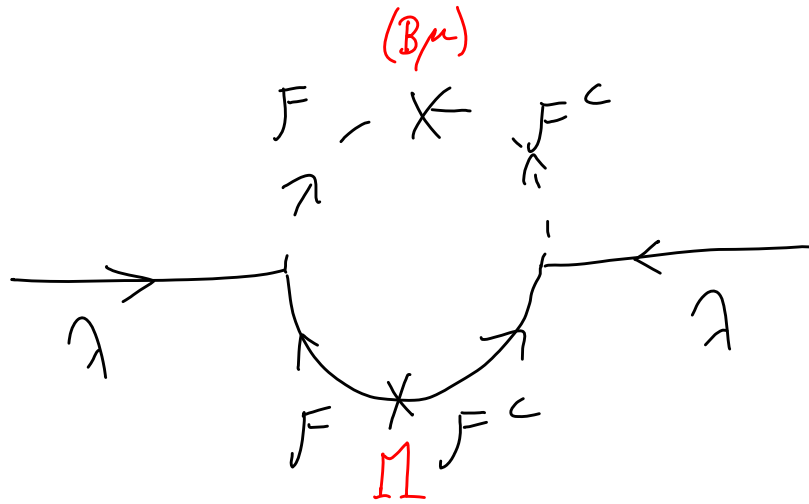
## G-M Vector

$$\int d^4\theta \phi^\dagger \phi [F^\dagger F + F^{\dagger c} F^c + \lambda F F^c + h.c.]$$

$\Downarrow$

$$\int d^4\theta F^\dagger F + F^{\dagger c} F^c + \left(\frac{\phi^\dagger}{\phi}\right) \lambda F F^c + h.c.$$

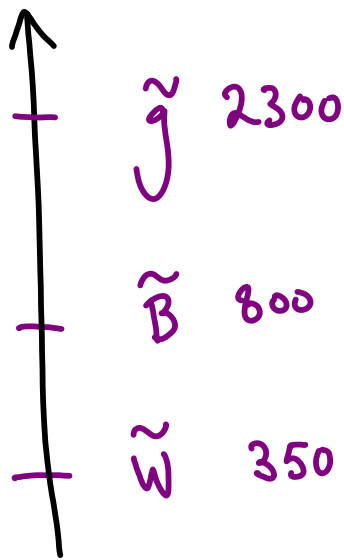
$\left(\frac{B\mu}{M}\right) = -m_{3/2}$ , off anom-med trajectory



$\left(\frac{B\mu}{M}\right) = m_{3/2}$  LE  $m_{\lambda}'s$  UV insensitive

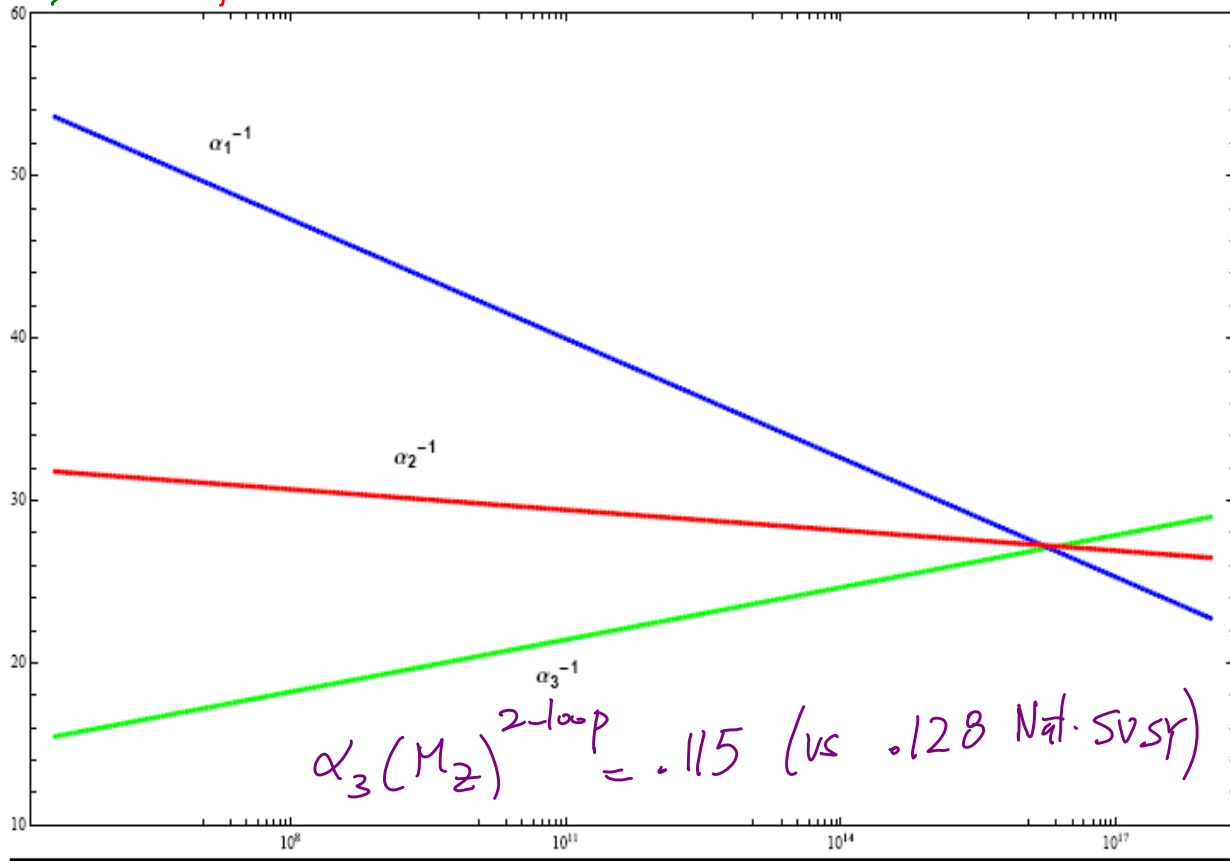
$\left(\frac{B\mu}{M}\right) \neq m_{3/2}$  Big Change to Anom-med Spectrum

# "Simplest" Split SUSY



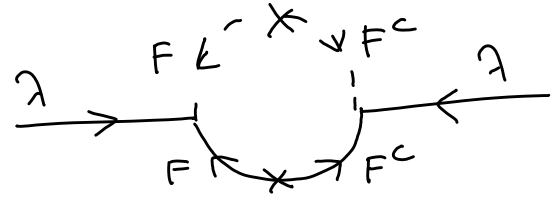
$$\mu = m_{3/2} = 100 \text{ TeV}$$
$$\left\{ \begin{array}{l} \tan\beta \sim 4, \\ m_H \sim 125 \text{ GeV} \end{array} \right\}$$

# Unification a Bit Better than Natural SUSY





# Additional G-M $(5 + \bar{5})'_s$



$\alpha_s(M_2)$   
 $= .106$   
 $-.107$

N	$\tan \beta$	$m_s$ (TeV)	$m_{3/2}$ (TeV)	$\mu$ (TeV)	$M_1$ (GeV)	$M_2$ (GeV)	$M_3$ (GeV)
0	4.0	80	80	80	802	377	2394
1	2.8	300	75	300	974	773	779
2	3.0	100	80	100	1307	1377	1478
3	2.7	250	50	50	945	1113	1951
4	2.7	250	50	50	1094	1409	2955
5	2.9	150	30	30	742	1025	2457

Can Get  
 Quite  
 Compressed  
 Spectra!

Values of  $\tan \beta$  and scalar mass  $m_s$  chosen to yield higgs mass of 125 GeV;  $m_{3/2}$  is gravitino mass, messenger scale  $M = \lambda m_{3/2}$ , GM term coupling  $\lambda$  chosen such that  $M = m_s$ .

More Optimistic for LHC Discovery

Some Comments on Flavor



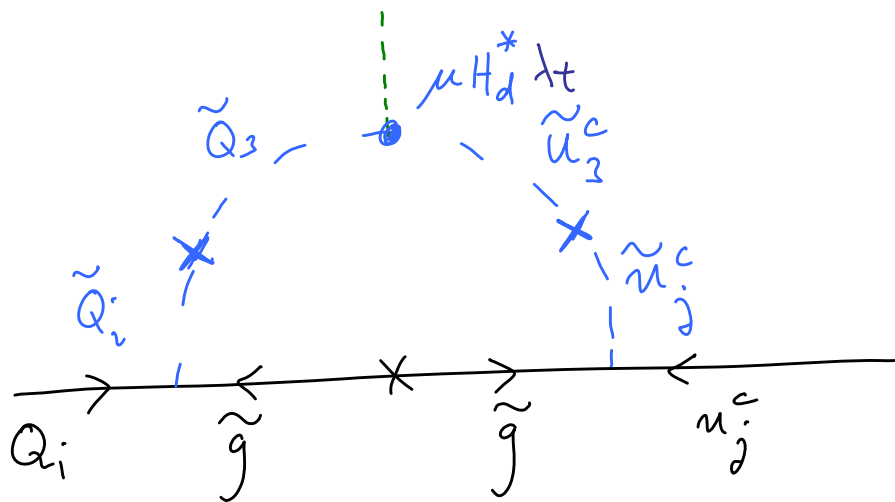
Naturalness' Loss = Flavor Gain

100-  
1000 TeV ↑  
10 TeV ↑  
1 TeV ↑

↑ Not KM like  
↑ Quasi KM-like  
↑ KM-like

can't predict what will first be seen!

Generic Soft Masses  $\rightarrow$  Need For Split Spectrum



Way too big  
unless split spectrum!



$$\delta \lambda_{up}^{ij} \sim \lambda_t \left( \frac{\alpha_s}{\pi} \right) \left[ \frac{\mu m \tilde{g}}{m_s^2} \right] \sim 10^{-2} \times \left[ \frac{\mu m \tilde{g}}{m_s^2} \right]$$

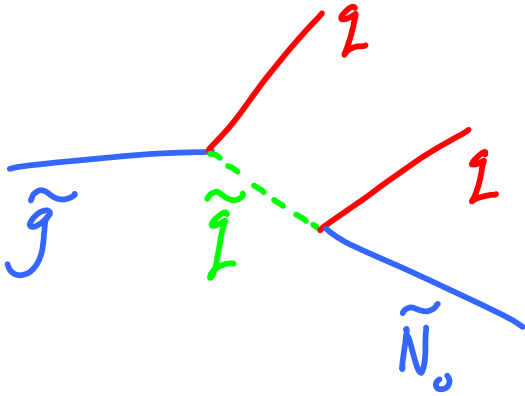
• With  $\mu \sim m_S$ ,  $m_{\text{gaugino}} \sim \alpha m_S$ ,

theories of radiative fermion masses  
are easy to build! SUSY was  
always crucial for this idea — naturalness  
made it incredibly hard — minimal split  
makes it easy!

LHC Signals, Smoking Guns

"How can we probe higher scales?"

Possible smoking gun for heavy scalars:



Long lived gluino:

visible displaced vertex

$$m_{\tilde{g}} \gtrsim \left( \frac{\Delta m}{1 \text{ TeV}} \right)^{5/4} \times (400 \text{ TeV})$$

Note that with scalars around  $\sim 100$ 's  
TeV, displacements in the  $100$ 's  $\mu\text{m}$ 's  
 $\rightarrow$  cm's are expected. Interesting  
Challenges to see them.

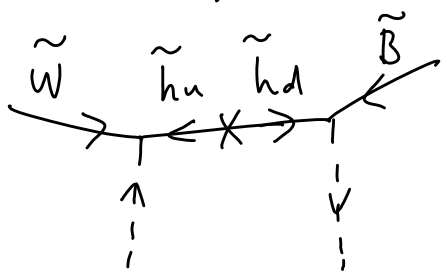


# Ewk Production

- With gauginos + higgsinos light, usual multilepton + MET searches
- Extremely constrained + simple story for simplest split:

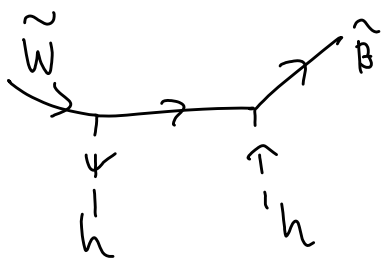
\* Only interactions between  $\tilde{W}, \tilde{b}$  from higher-

dim ops:



$$\frac{g^2 (\tilde{h}^\dagger \tilde{W} \tilde{h}) \tilde{B}}{M}$$

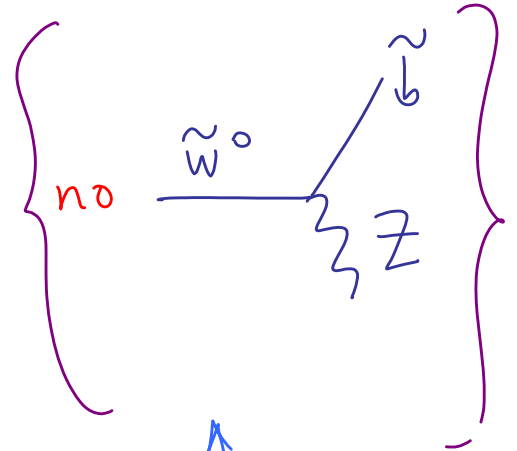
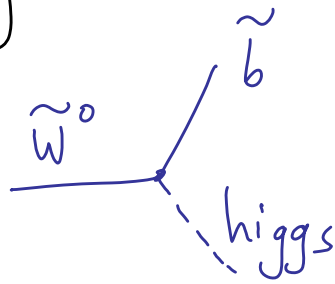
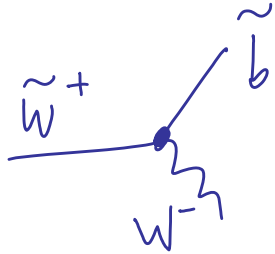
dim 5



$$\frac{g^2 (\tilde{h}^\dagger \tilde{D} \tilde{h}) \tilde{W} \tilde{\sigma}_\mu \tilde{B}}{M^2}$$

dim 6

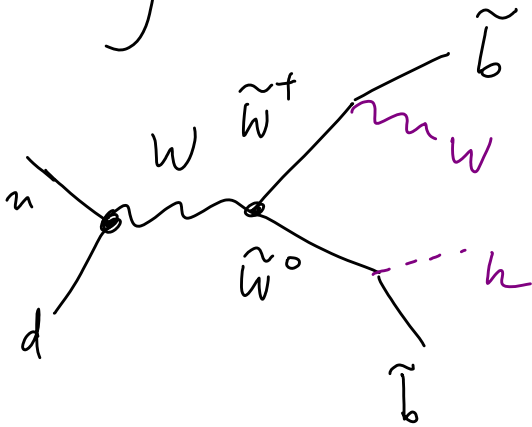
# Leading couplings



[ $\tilde{W}^+ / \tilde{W}^0$  nearly degenerate]

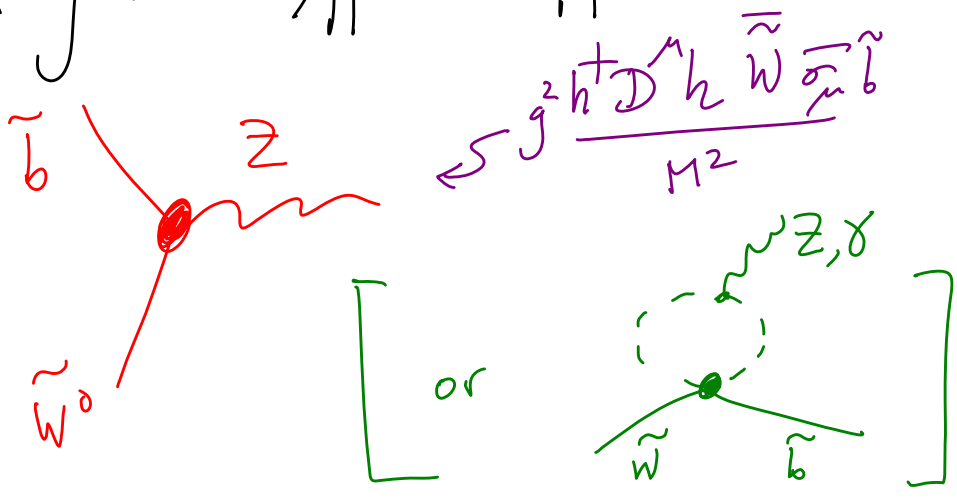
Fingerprint of Heavy Higgsinos!

Only direct EWK production:



Measuring tiny Br for  $\tilde{b} \rightarrow \tilde{W}^0 Z$

(or  $\tilde{W}^0 \rightarrow \tilde{b} Z$ ) is direct probe of  $\text{dim } 6$  op,  
 which goes as  $\frac{1}{M^2}$  not  $\frac{1}{M^4}$ !



$$\leftarrow \frac{g^2 h^+ D^{\mu} h \tilde{W}^{\mu} \tilde{b}}{M^2}$$

$$\text{Br} [\tilde{W}^0 \rightarrow \tilde{b} Z] \sim \left( \frac{m_Z}{M} \right)^2$$

Probes heavy Higgsino, directly tuning

What about

$$g^2 \frac{(\tilde{h}^\dagger \tilde{h}_{nu})(h h_d)}{M} \rightarrow \text{Split } \tilde{h}^+ \tilde{h}^0$$

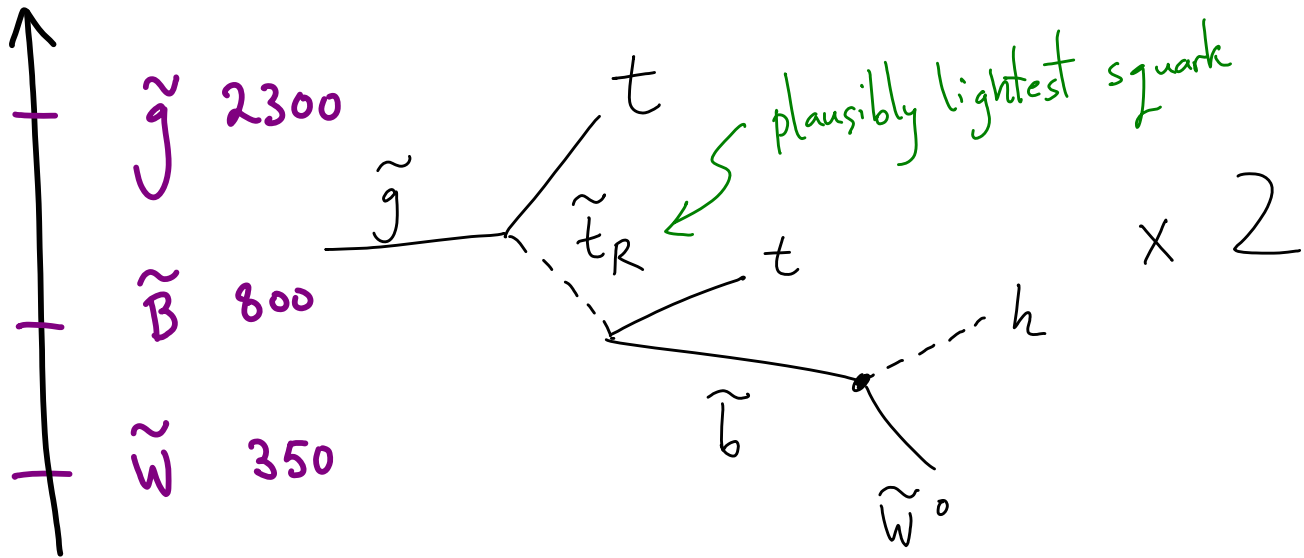
$$\begin{array}{l} \text{=====} \\ \text{=====} \\ \text{=====} \end{array} \tilde{w}, \tilde{b}$$

$$\text{-----} \tilde{h}_{u,d}$$

$$\Delta m \sim \left[ \alpha M_2 + c \frac{M_2^2}{M} \right]$$

$$\Gamma \propto \Delta m^5, \quad \Gamma^{-1} \sim 10 \text{ cm!}$$

$$\text{Derivation} \Rightarrow M_{\tilde{w}, \tilde{b}} \sim \alpha^{-1} M_{\tilde{h}} \text{ (wavy red line)}$$



$\Rightarrow$  8  $b$ 's, 4  $W$ 's [ + perhaps displacement ]!  
 in every event!



Finding any deviation in  $\sigma \times Br$   
for the Higgs, would instantly  
demolish this picture.

As many people have remarked,  
 $m_H \sim 125$  seems almost  
maliciously designed to prolong the  
agonny of BSM theorists....

# HIGH DRAMA FOR 2012



- ★ Higgs  $\sigma \times \text{Br}$  [+ confirmation !!]
- ★ Stops
- ★ Pushing Gluino limits to 1 TeV in  $SS$  leptons  
+  $b^r$ 's searches
- ★ Multileptons
- .....

# The Stakes Are Very High

$M_H \sim 125 \text{ GeV}$

11<sup>th</sup> hour  
naturalness  
(remember  
COBE!)

Somewhat  
elaborate

Un-natural

Simple

(Even minimal  
split is  
dramatic  
tuning!)

